Calculations for Nondestructive BEC Imaging

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1 Diffraction by a Phase Object

We describe the phase object (BEC) by the deviation $\delta n(x, y, z)$ of the local index n(x, y, z) from unity:

$$\delta n(x, y, z) := n(x, y, z) - 1.$$
 (1)

In the thin-lens approximation, we can integrate over the propagation direction (z) of the probe, and consider only the 2-dimensional index profile:

$$\delta n_z(x,y) := \int \delta n(x,y,z) \, dz. \tag{2}$$

This phase shift imparts momentum to the light as follows: in the scalar-wave approximation, the mean-square transverse wave vector $\langle k_{\rm T}^2 \rangle$ due to the relative phase shift of $\exp[ik\delta n_z(x,y)]$ (k is the total wave vector) is

$$\langle k_{\rm T}^2 \rangle = \frac{1}{|E|^2} \int E^*(x,y) (k_x^2 + k_y^2) E(x,y) \, dx \, dy$$

$$= -\int e^{-ik\delta n_z} (\partial_x^2 + \partial_y^2) e^{ik\delta n_z} \, dx \, dy$$

$$= k^2 \int \left[(\partial_x \delta n_z)^2 + (\partial_y \delta n_z)^2 \right] \, dx \, dy.$$

$$(3)$$

Here we have dropped terms of the form $\partial_x^2 \delta n_z$, which do not contribute if we assume that δn_z is purely real (i.e., we ignore absorption effects).

To make this expression more concrete, we can consider a Gaussian approximation to the true BEC atomic distribution,

$$\delta n_z = \frac{\delta \phi_a}{2\pi\sigma_x \sigma_y} \exp\left[-\frac{x^2}{2\sigma_x^2} - \frac{y^2}{2\sigma_y^2}\right],\tag{4}$$

where $\delta \phi_{a}$ is the integrated phase shift:

$$\delta\phi_{\mathbf{a}} := \int \delta n(x, y, z) \, dx \, dy \, dz. \tag{5}$$

Then we can write

$$\langle k_{\rm T}^2 \rangle = \frac{(\delta\phi_{\rm a})^2 (\sigma_x^2 + \sigma_y^2) k^2}{8\pi \sigma_x^3 \sigma_y^3} \tag{6}$$

for the optical transverse momentum in the Gaussian approximation.

2 ATOMIC PARAMETERS

This momentum is similarly imparted to the atomic cloud, and the atomic momentum diffusion is given by $\langle \hbar^2 k_T^2 \rangle$ multiplied by the incident photon flux. Thus we can write

$$D = \frac{I\hbar}{N\omega} \langle k_{\rm T}^2 \rangle \tag{7}$$

for the momentum diffusion rate per atom, where I is the probe intensity, ω is the optical frequency, and N is the number of atoms in the condensate.

2 Atomic Parameters

We can write the local index as (see, for example, Cohen-Tannoudji, Dupont-Roc, and Grynberg, *Atom-Photon Interactions*, p. 604)

$$\delta n = -\frac{n_{\rm d}|d|^2}{2\epsilon_0 \hbar} \frac{\Delta}{\Delta^2 + \Gamma^2/4 + \Omega^2/2},\tag{8}$$

where n_d is the local number density, d is the appropriate dipole moment (here, approximately the effective dipole moment for far-detuned, linearly polarized light, see "Rubidium 87 D Line Data" for details). In the far-detuned limit, the integrated phase shift is then

$$\delta\phi_{\rm a} = -\frac{N|d|^2}{2\epsilon_0 \hbar\Delta}.\tag{9}$$

Thus the diffusion rate in the Gaussian approximation is

$$D = (\hbar^2 k^2) \frac{N|d|^4}{4\epsilon_0^2 \hbar^3 \omega} \frac{I}{\Delta^2} \frac{\sigma_x^2 + \sigma_y^2}{8\pi \sigma_x^3 \sigma_y^3},$$
(10)

which has an overall scaling similar to spontaneous scattering, I/Δ^2 .

3 Signal/Noise

From the Kadlecek et al. paper, the signal/noise ratio is given by

$$S/N = \phi \sqrt{2\eta N_{\rm p}},\tag{11}$$

where $N_{\rm p}$ is the number of photons striking a particular CCD pixel, η is the CCD quantum efficiency, we consider only the limit of large reference-beam intensity, and we have reduced the value quoted here by a factor of $\sqrt{2}$ from the value quoted in the paper, as is appropriate for spatial-heterodyne mode (since we must average over the heterodyne phase in the mean-square sense). This expression is also valid only for small local phase shifts ϕ ; for large ϕ , we must average also over ϕ , and in this case the signal/noise is given by the same expression, but with ϕ replaced by 1. To evaluate the signal/noise ratio, we simply note that the typical phase shift is

$$\phi \approx -\frac{N|d|^2k}{2\epsilon_0 \hbar \Delta \sigma_x \sigma_y},\tag{12}$$

and the photon number is

$$N_{\rm p} = \frac{I t_{\rm exp} A_{\rm pix}}{\hbar \omega},\tag{13}$$

where t_{exp} is the exposure time of the image, and A_{pix} is the effective CCD pixel area, taking into account any magnification in the imaging system. Thus in the small ϕ regime, where the experiment is likely to operate, the signal/noise ratio scales as \sqrt{I}/Δ^2 .

 $\begin{aligned} H_{Z} &:= \frac{1}{\sec} \qquad M_{H_{Z}} := 10^{6} \cdot H_{Z} \qquad k_{H_{Z}} := 10^{3} \cdot H_{Z} \qquad G_{H_{Z}} := 10^{9} \cdot H_{Z} \qquad T_{H_{Z}} := 10^{12} \cdot H_{Z} \\ m_{W} &:= 10^{-3} \cdot watt \qquad m_{S} := 10^{-3} \cdot \sec \qquad c_{m} := 10^{-2} \cdot m \\ h_{D}ar &:= 1.054571596 \cdot 10^{-34} \cdot joule \cdot \sec \qquad \epsilon_{0} := 8.854187817 \cdot 10^{-12} \cdot \frac{f_{arad}}{m} \qquad \mu_{m} := 10^{-6} \cdot m \end{aligned}$

Parameters for Rb 87

$$c := 2.99792458 \cdot 10^{8} \cdot \frac{m}{\text{sec}} \qquad \omega := 2 \cdot \pi \cdot 384.2279818773 \cdot \text{THz} \qquad \lambda := 2 \cdot \pi \cdot \frac{c}{\omega} \qquad k := \frac{2 \cdot \pi}{\lambda}$$

$$\Gamma := 2 \cdot \pi \cdot 6.065 \cdot \text{MHz} \qquad d := 2.069 \cdot 10^{-29} \cdot \text{coul} \cdot \text{m}$$

$$N := 2 \cdot 10^{5} \qquad \text{ox} := 100 \cdot \mu \text{m} \qquad \text{oy} := \text{ox} \qquad \text{Isat} := \frac{c \cdot \epsilon 0 \cdot \Gamma^{2} \cdot \text{hbar}^{2}}{4 \cdot d^{2}} \qquad \text{Isat} = 2.504 \cdot \frac{\text{mW}}{\text{cm}^{2}}$$

$$\Delta := 2 \cdot \pi \cdot 2 \cdot \text{GHz} \qquad \Gamma := 2 \cdot \frac{\text{mW}}{\text{cm}^{2}} \qquad \delta \phi a = -3.64826 \cdot \mu \text{m}^{3} \qquad 4 \cdot \left(\frac{\Delta}{\Gamma}\right)^{2} = 4.35 \cdot 10^{5}$$

$$\frac{\Gamma}{\text{Isat}} = 0.799$$

Diffusion due to coherent lensing

$$D := \frac{N \cdot d^4}{4 \cdot \varepsilon 0^2 \cdot hbar^3 \cdot \omega} \cdot \frac{I}{\Delta^2} \cdot \frac{\sigma x^2 + \sigma y^2}{8 \cdot \pi \cdot \sigma x^3 \cdot \sigma y^3} \cdot hbar^2 \cdot k^2$$
$$D = 4.16 \cdot 10^{-9} \cdot \frac{hbar^2 \cdot k^2}{ms}$$

Diffusion due to spontaneous emission

$$\operatorname{Rsc} := \left(\frac{\Gamma}{2}\right) \cdot \left(\frac{I}{\operatorname{Isat}}\right) \cdot \left[1 + 4 \cdot \left(\frac{\Delta}{\Gamma}\right)^2 + \left(\frac{I}{\operatorname{Isat}}\right)\right]^{-1} \qquad \operatorname{Rsc} = 0.035 \cdot \frac{1}{\operatorname{ms}}$$
$$\operatorname{Dsp} := \operatorname{Rsc} \cdot \operatorname{hbar}^2 \cdot \operatorname{k}^2 \qquad \qquad \operatorname{Dsp} = 0.035 \cdot \frac{\operatorname{hbar}^2 \cdot \operatorname{k}^2}{\operatorname{ms}}$$

Signal/noise ratio

$$\phi := -\frac{N \cdot d^2 \cdot k}{2 \cdot \epsilon 0 \cdot hbar \cdot \Delta \cdot \sigma x \cdot \sigma y} \qquad \phi = -0.003 \qquad Apix := (10 \cdot \mu m)^2$$
$$\eta := 0.3 \qquad Np := \frac{I \cdot Apix}{hbar \cdot \omega} \qquad Np = 7.856 \cdot 10^6 \cdot \frac{1}{ms}$$
$$SN1 := |\phi| \cdot \sqrt{2 \cdot \eta \cdot Np} \qquad SN1 = 6.378 \cdot \frac{1}{\sqrt{ms}} \qquad (valid for small phase shift)$$